

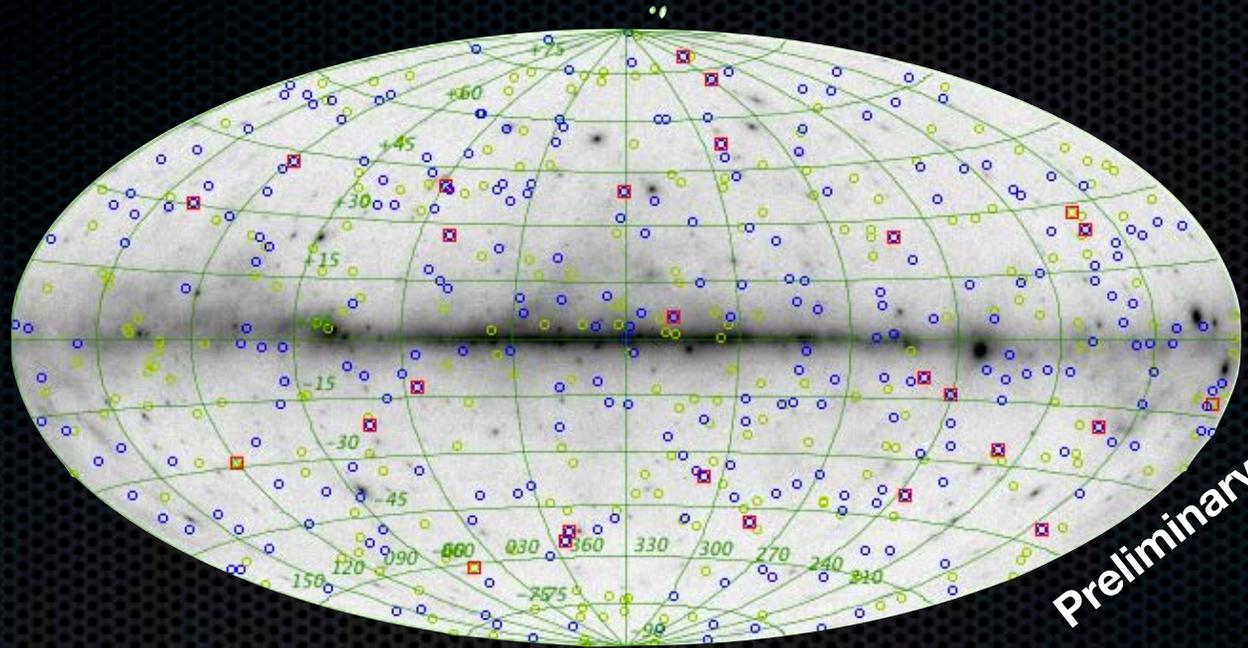
On The Lack of LAT Detected GRBs

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On behalf of the Fermi collaboration

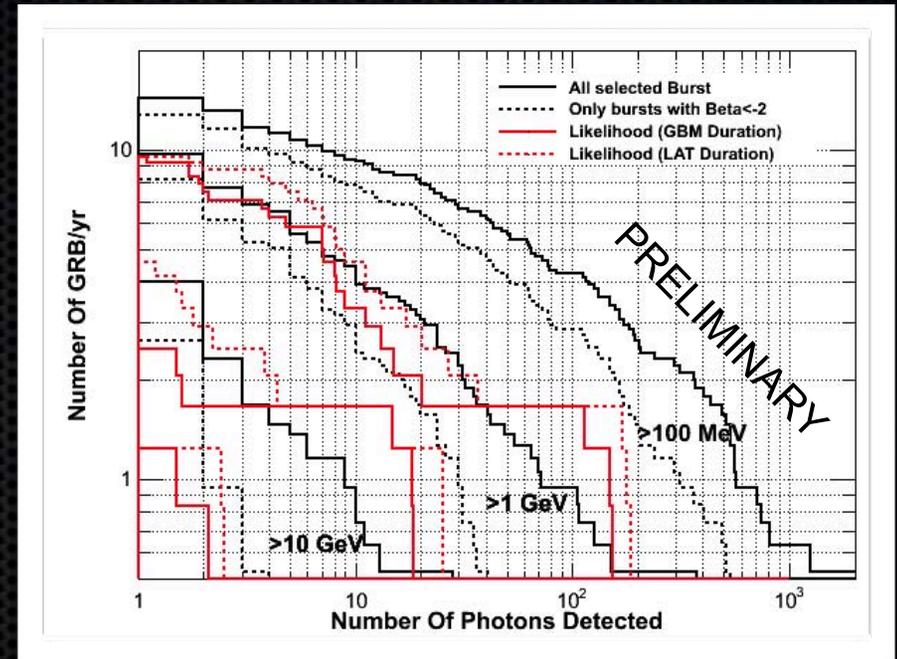
Fermi GRB Detections



- ✦ GBM Detected GRBs (until March 1st): **620** - Blue
- ✦ GRBs in LAT FOV: **288** (46%) - Green
- ✦ LAT Detected GRBs (>100 MeV): **23** (8%) - Red
- ✦ LAT LLE Only Detected GRBs: **5** (2%)

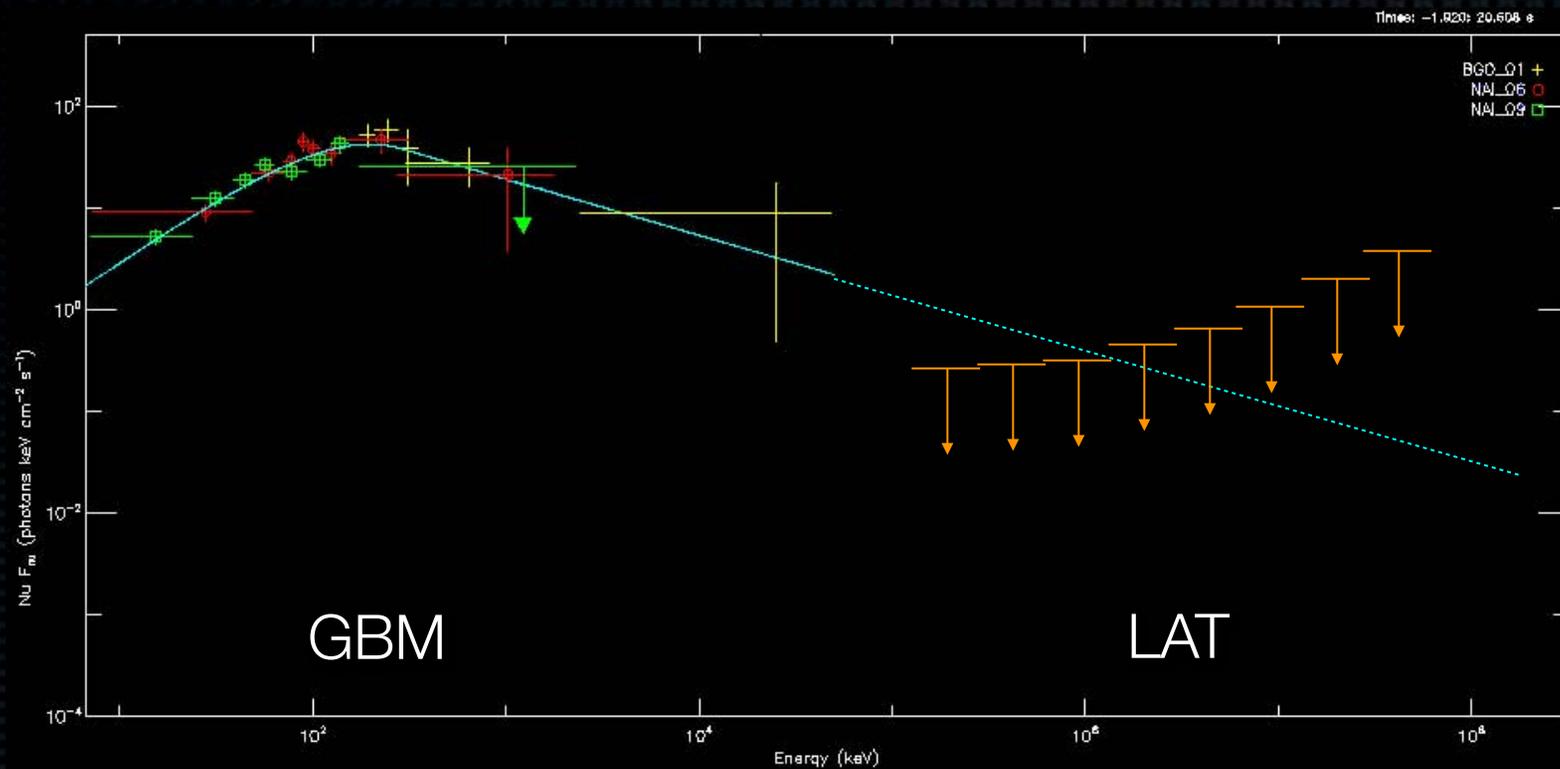
Expected Detection Rate

- Take BATSE spectra, extrapolate and compare to actual detection rate
 - Predicted: 9.3 GRBs/year > 100 MeV
 - Observed: 8.0 GRBs/year > 100 MeV
- This includes GRBs with extra components
- We are seeing fewer GRBs than predicted, especially at GeV energies
- Possible explanations
 - High energy emission is suppressed
 - Extrapolations are uncertain
- Extra components must be rare!



Omodei's Presentation

Spectral Fits

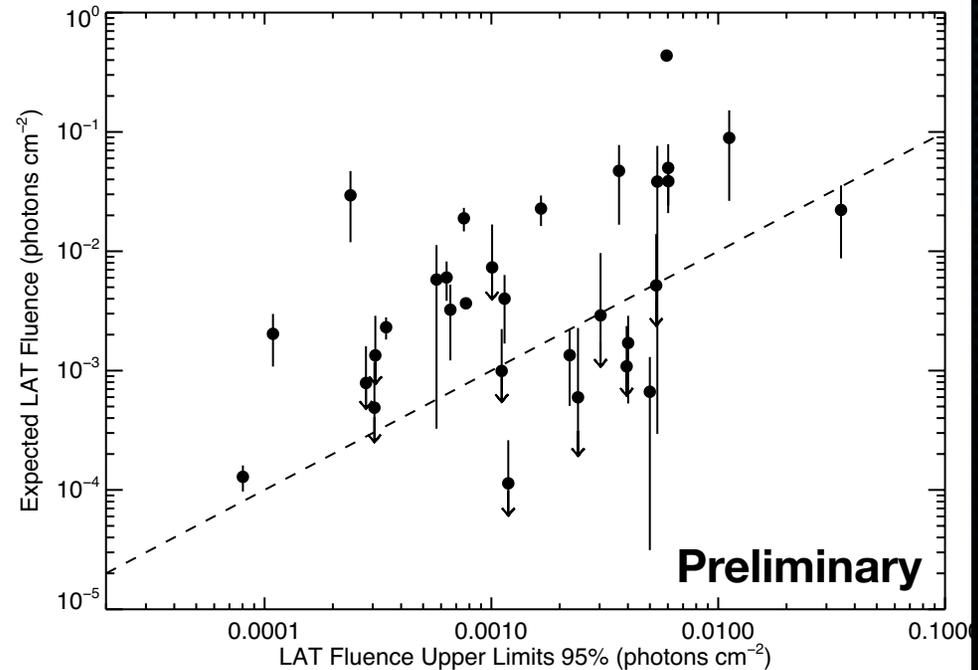
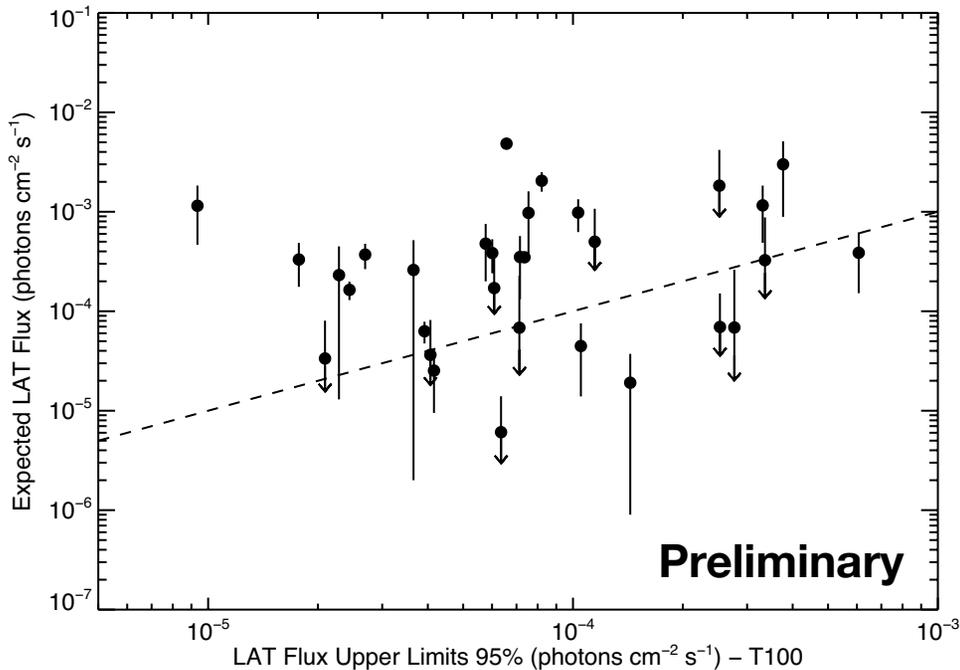


- ✦ Fit NaI+BGO spectrum from 8 keV to 40 MeV in RMFIT
- ✦ Estimate the expected flux in the 100 MeV to 10 GeV range
- ✦ Compare upper limits to the expected LAT flux

Spectroscopic Sample

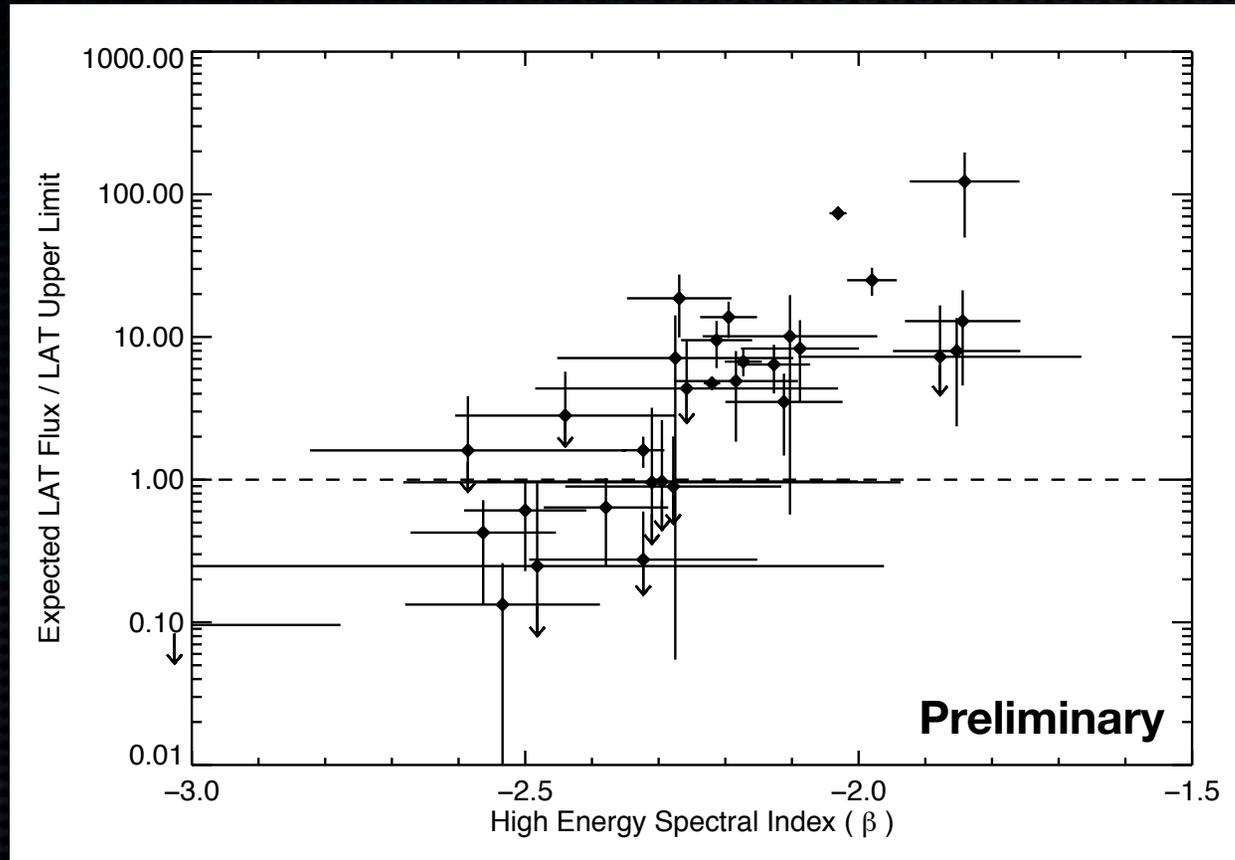
- Bright BGO Sample:
 - GRBs with 70 cts/s in BGO in LAT FOV: 92
 - “LAT Dark GRBs” (i.e. no LAT detection)
- “Gold” Sample:
 - Number of bright BGO GRBs with $\Delta\text{Beta} < 0.5$: 30
- Expected LAT Flux
 - Extrapolate β to find expected LAT flux
 - We use the full covariance matrix to estimate beta error

Expected Flux Comparisons



- ✦ 15 of the 30 GRBs have expected photon flux that exceed the T90 LAT photon flux upper limit
- ✦ Same for the expected photon fluence and LAT fluence upper limit

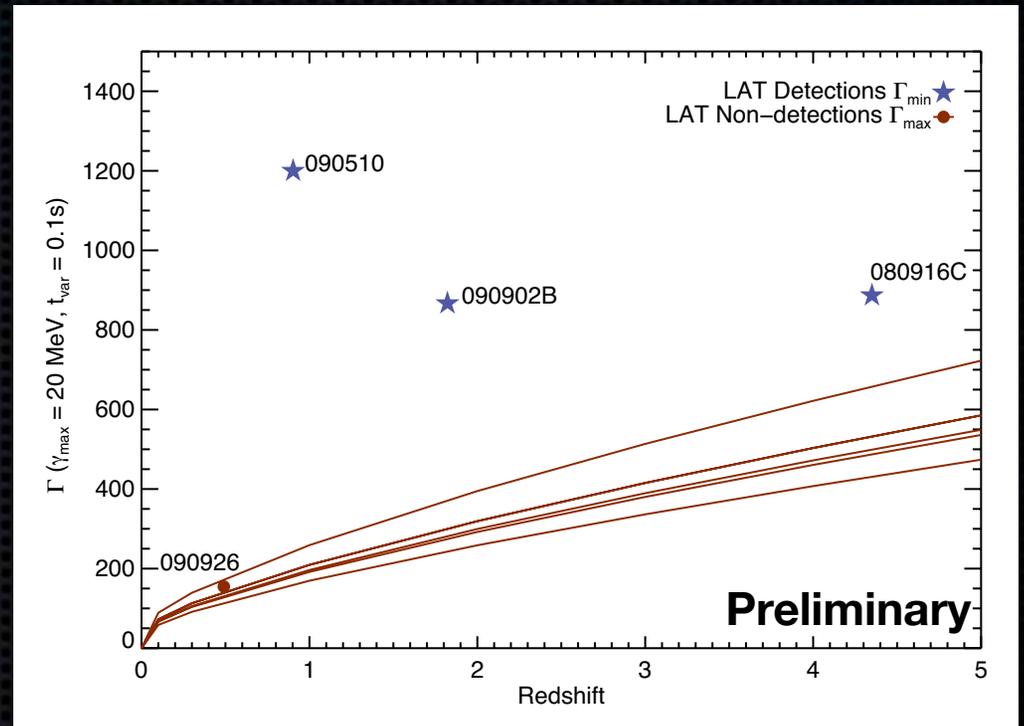
Beta vs Ratio



- GRBs with values of $\beta > -2.2$ typically exceed the LAT upper limits

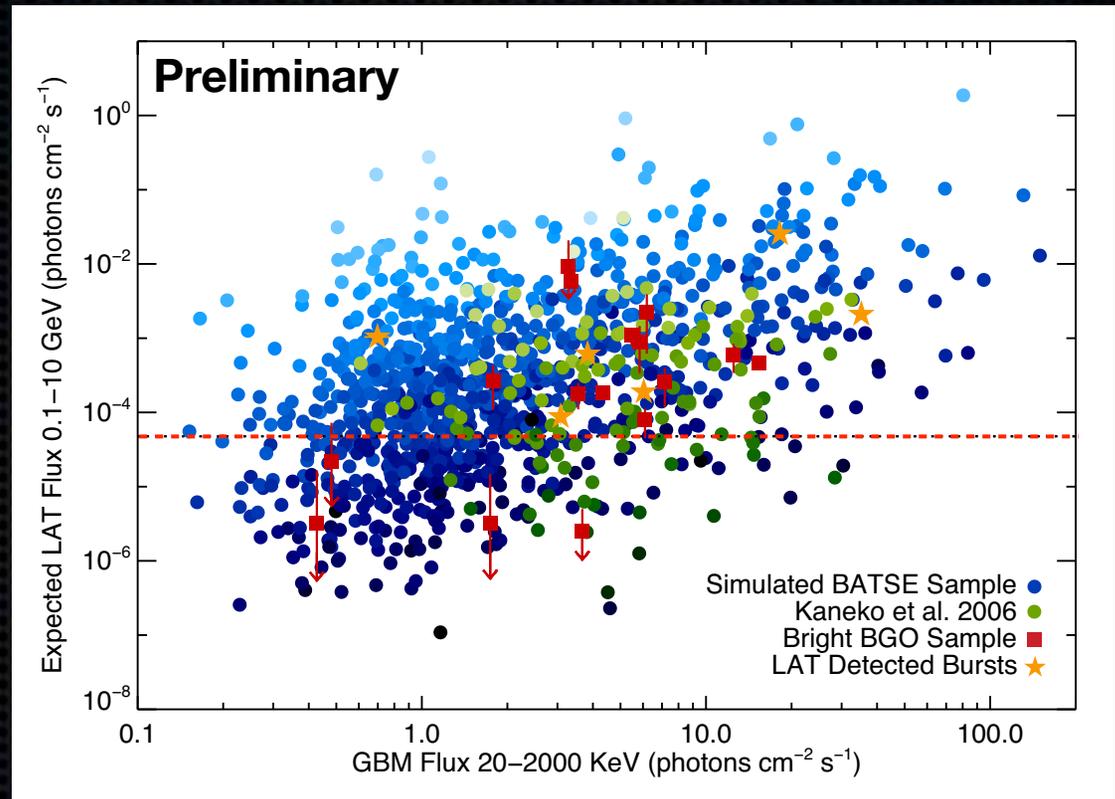
Lorentz Factor Distribution

- 3 LAT detected bursts have $\Gamma_{\min} > 800$
- For 6 LAT dark GRBs:
 - $\Delta t \sim 0.01\text{s}$ and $1 < z < 5$
 - If we assume $E_c \sim 100\text{ MeV}$
 - $\Gamma_{\max} \sim 50\text{-}600$
- LAT bursts may represent the high end of the Γ distribution
- LAT dark bursts may represent the low end of the Γ distribution



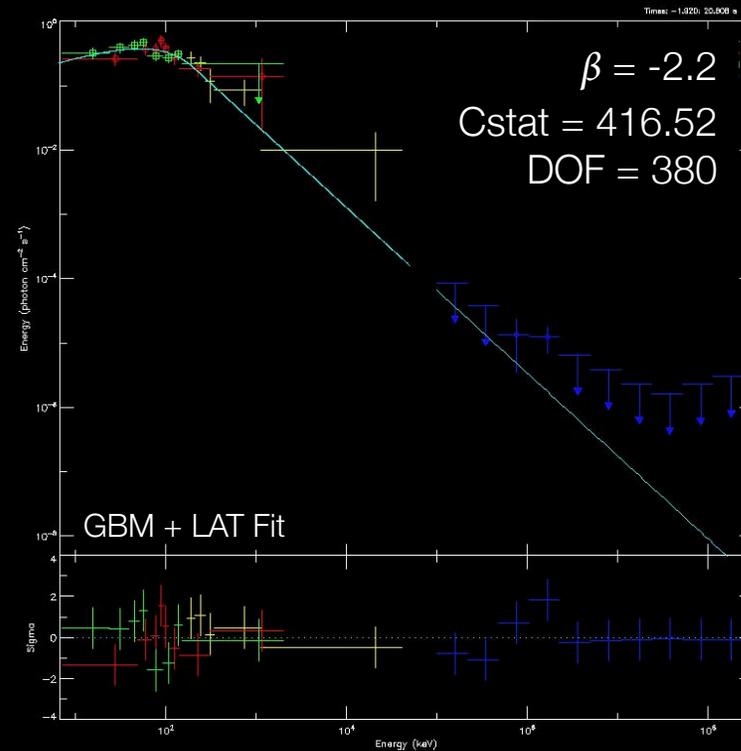
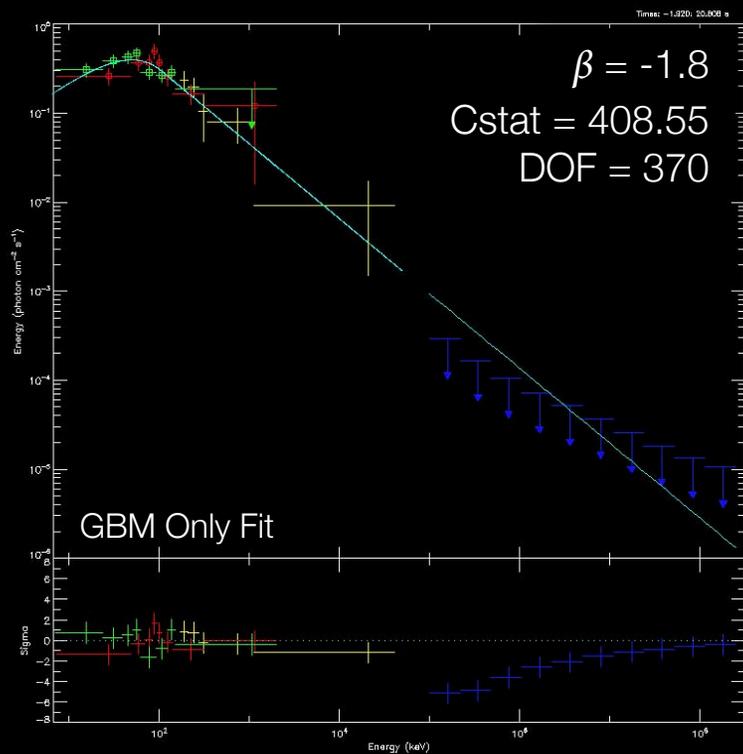
Expected LAT Flux

- ✦ Simulate a population of GRBs using BATSE E_{pk} , α , and β distributions
- ✦ Roughly 65-75% of a simulated BATSE sample have expected flux values that exceed the median 30s LAT sensitivity



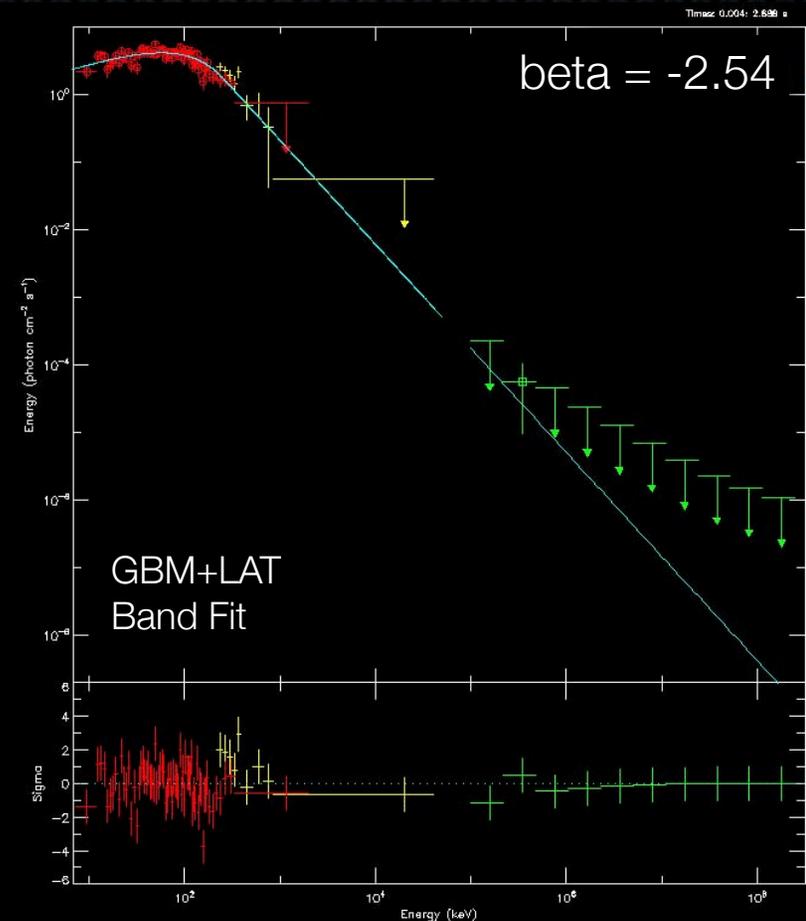
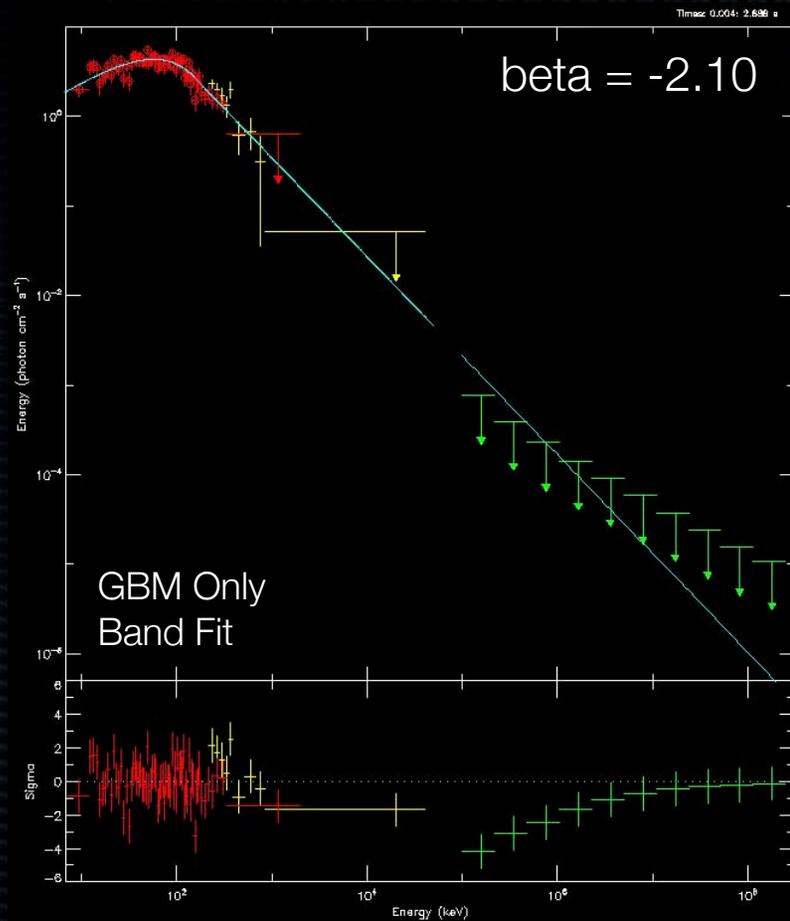
- ✦ High energy extrapolations must be misleading in order to explain the number of LAT “dark” bursts

Joint GBM+LAT Spectral Fits



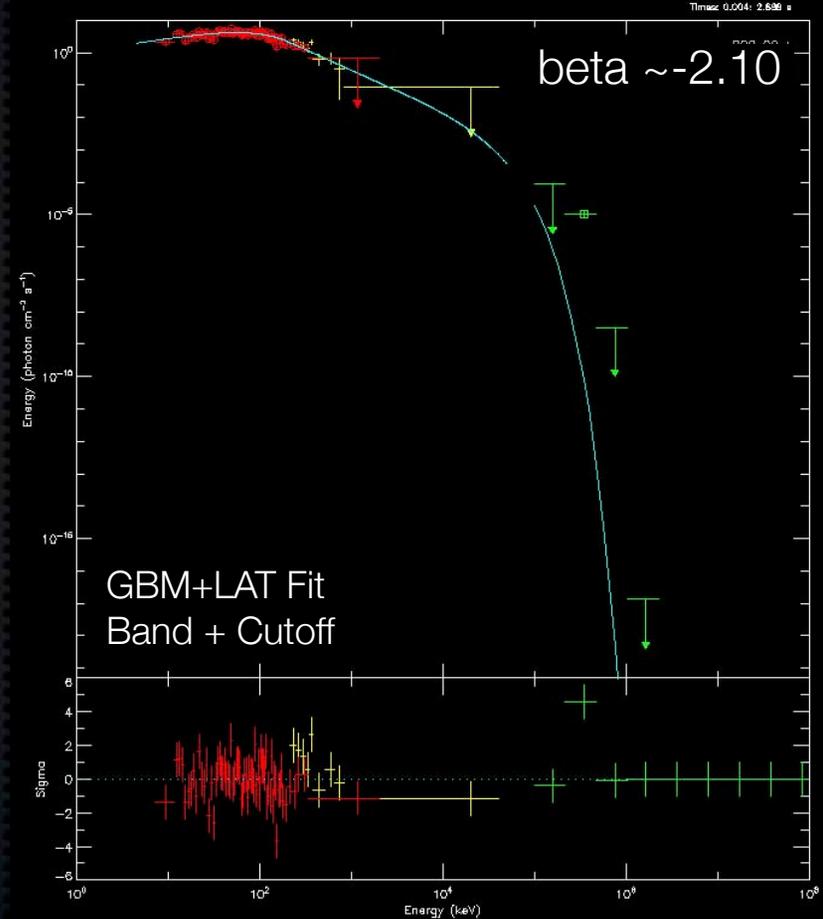
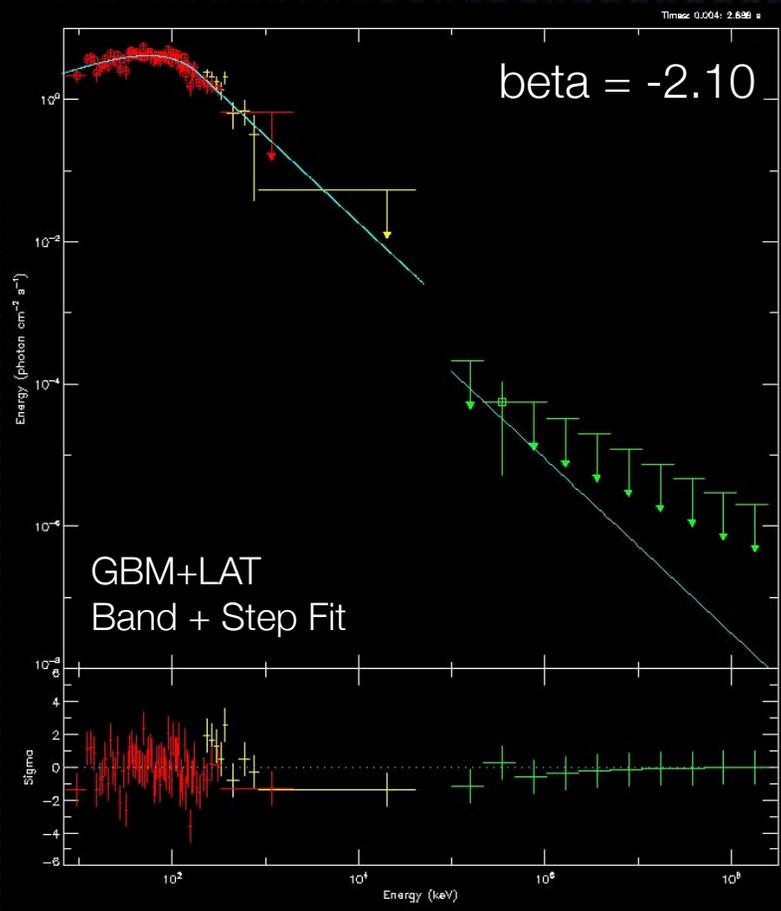
- ❖ Very different beta value if we include LAT limits in the spectral fits.
- ❖ For bright BGO sample, median $\beta = -2.2 \rightarrow -2.5$
- ❖ Which fit is statistically preferred?

Model Comparisons?



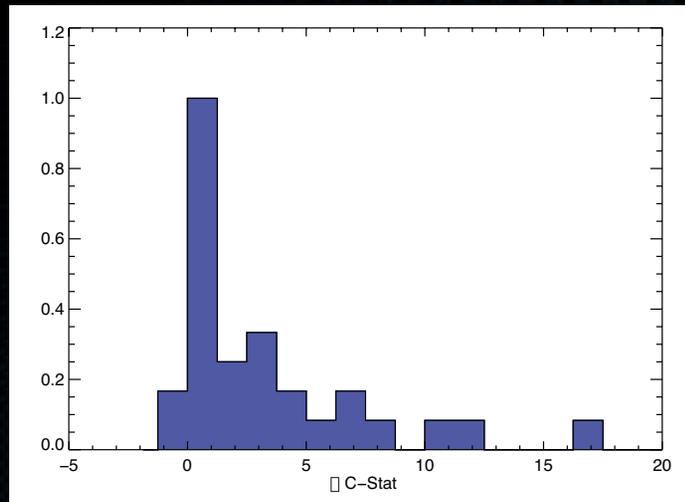
We cannot statistically compare these two scenarios using $\Delta\text{C-Stat}$ because we are using different data sets for the two fits

Nested Model Comparisons

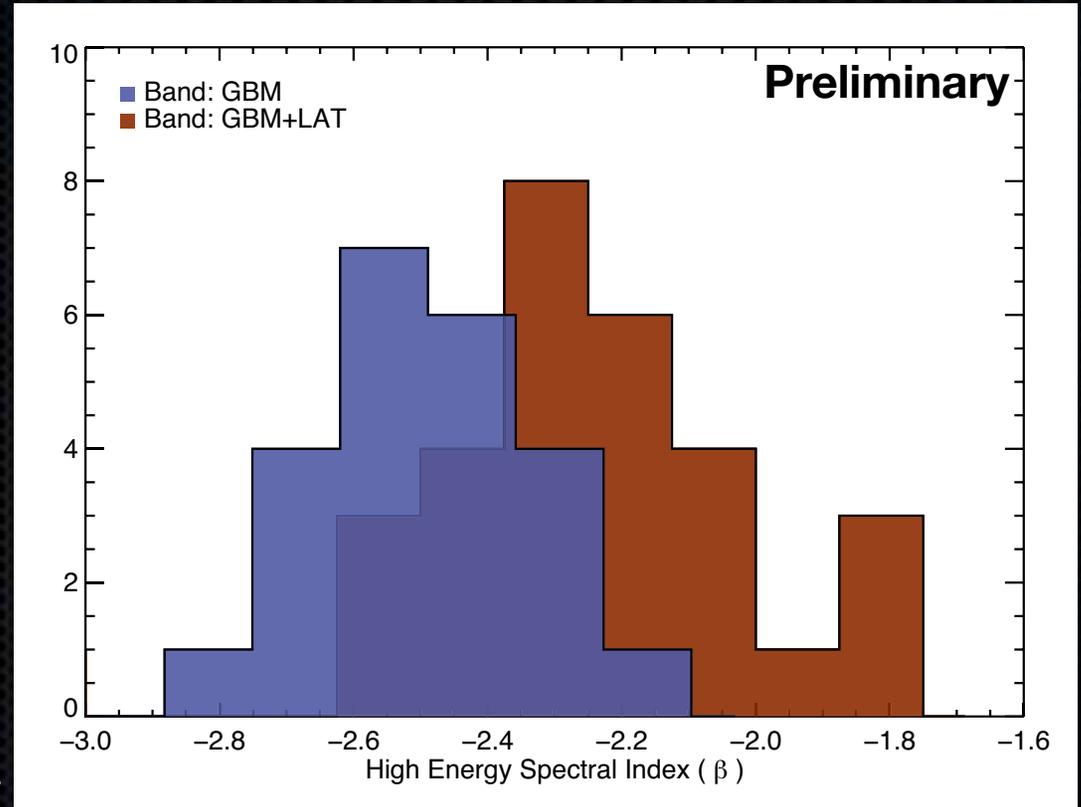


We have to compare the ΔC -stat values for the Band only, Band + Step Function, and Band+Cutoff fits to the same GBM+LAT data

ΔC -Stat

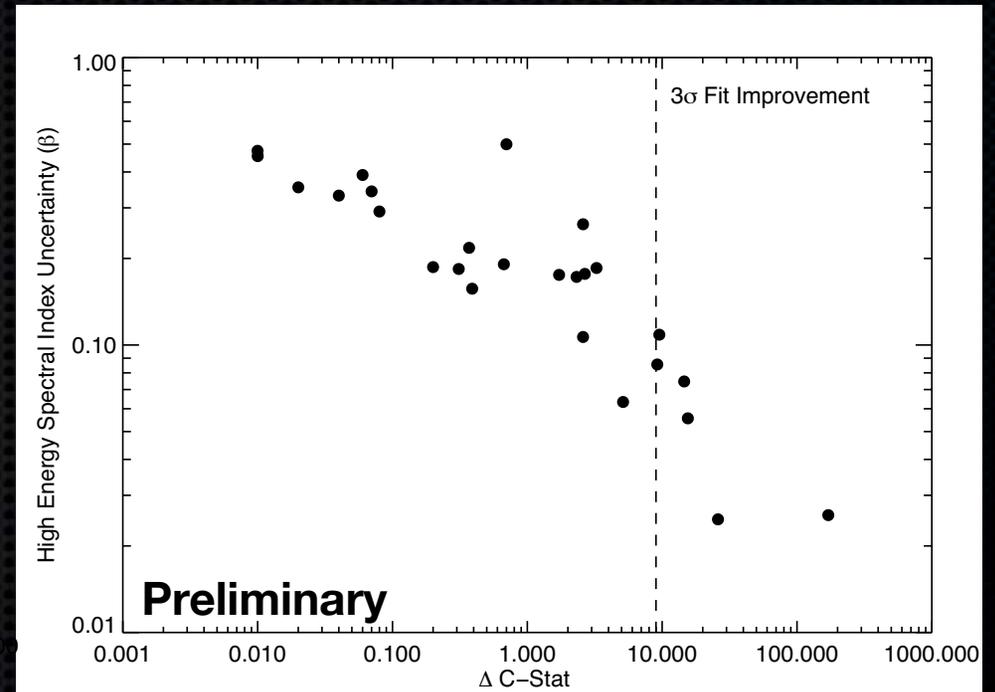
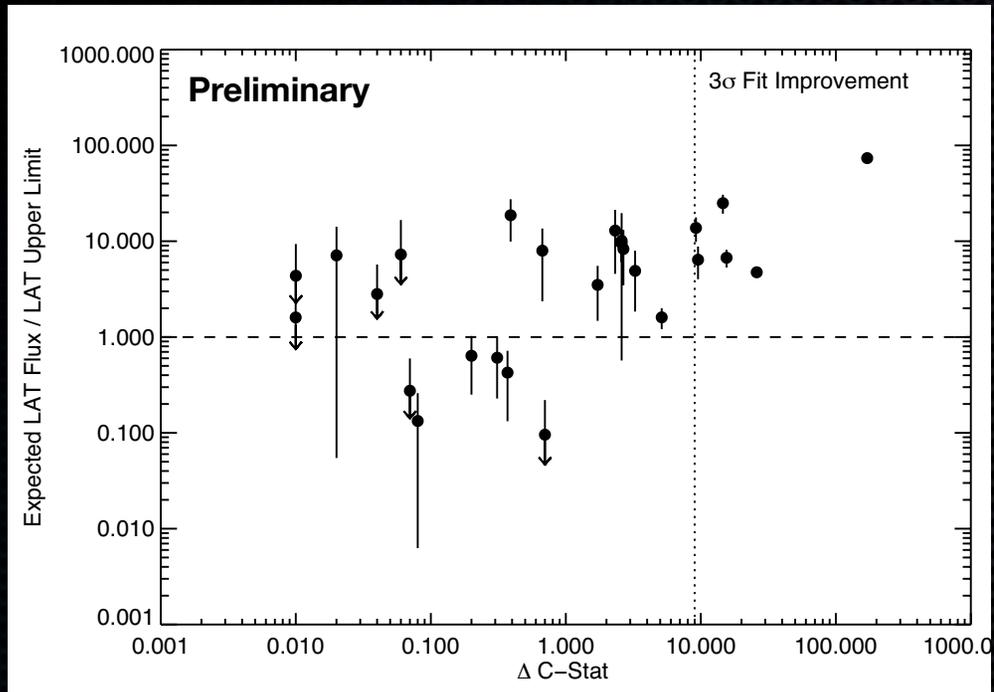


- Band vs. Band + Step Function, change of 1 degree of freedom
- Only 6 of 30 (20%) GRBs result in ΔC -Stat > 10
- We can reject the null hypothesis (the Band model) only for these bursts



- Band fit to GBM+LAT data results in softer beta values compared to fits to GBM data alone
- Possibly consistent with suggestions by Hascoet (this conference)

ΔC -Stat Correlations



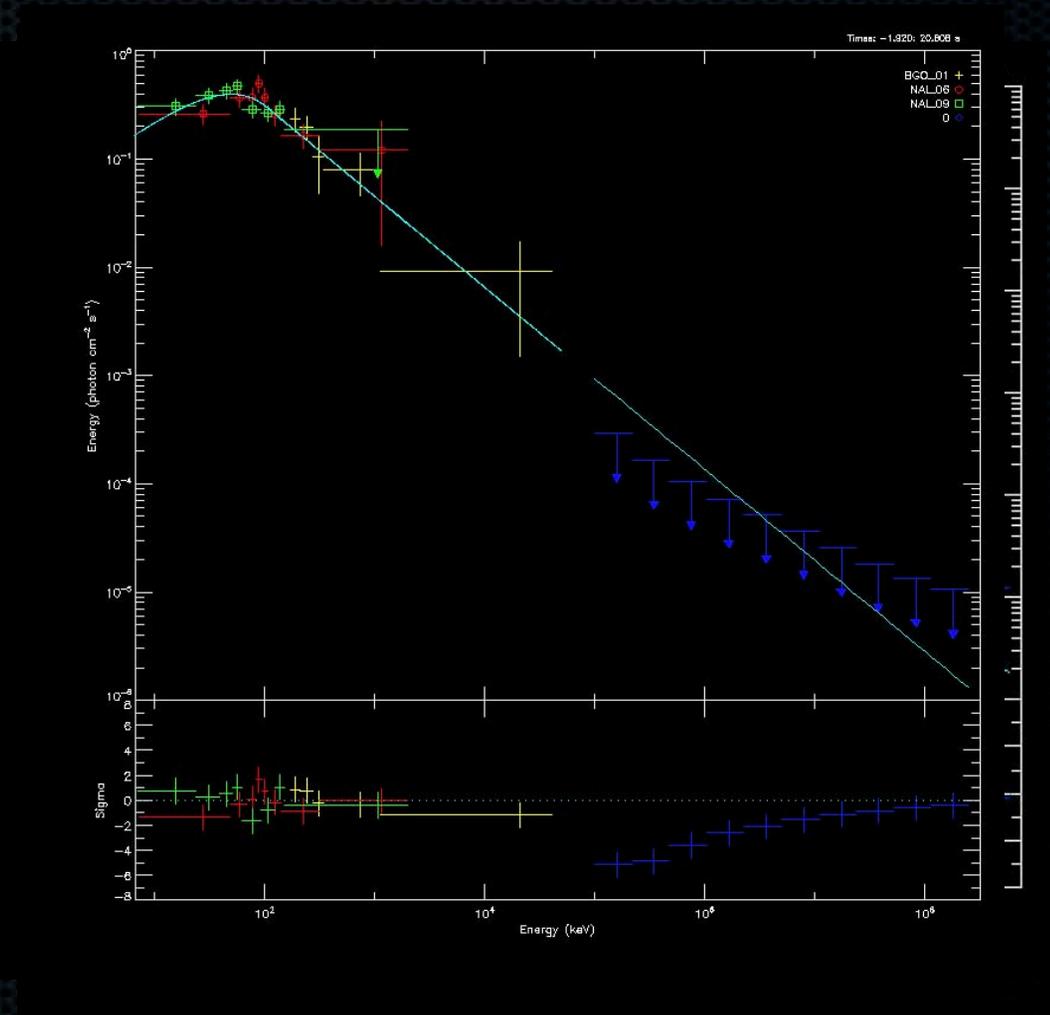
- Correlation between over-prediction of LAT flux and ΔC -Stat. Likewise, anti-correlation between σ_β and ΔC -Stat
- Statistical errors on β do not reflect the true, systematic, uncertainty in the parameter estimation

Conclusions

- GBM to LAT extrapolations can be misleading!
- Statistical uncertainties may not fully reflect the systematic uncertainties and cross-correlations among the spectral parameters
- ΔC -stat for a nested model comparison is the proper method of distinguishing between fits of increasingly complexity
- 24 (80%) GRBs in our spectroscopic sample are consistent with having a steeper beta value
- 6 of 30 (20%) prefer a spectral break
 - Two of these bursts show this break in the LLE selection
- Our previous estimates of the β distribution may be biased
- Use of future LLE data may help distinguish between cutoffs and softer β

Simulation Tests

- GRB 101113483
 - GBM Only: $\beta = -1.8$
 - GBM+LAT: $\beta = -2.2$
 - GBM+LAT+Step: $\beta = -1.8$
 - $\Delta C\text{-stat} \sim 5$
- Simulated GRB: $\beta = -1.8$
 - GBM Only: $\beta \sim -1.8$
 - GBM+LAT: $\beta \sim -2.2$
 - GBM+LAT+Step: $\beta = -1.8$
 - $\Delta C\text{-stat} \sim 25$



- Nested model comparison can distinguish the difference between the two scenarios, even though the different beta values are statistically excluded